

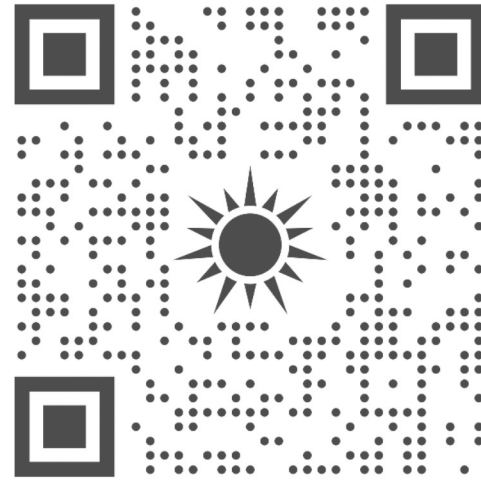
Dr. B Panos



Flare prediction with explainable artificial intelligence

SCOSTEP

16 May 2023



Prof. L Kleint



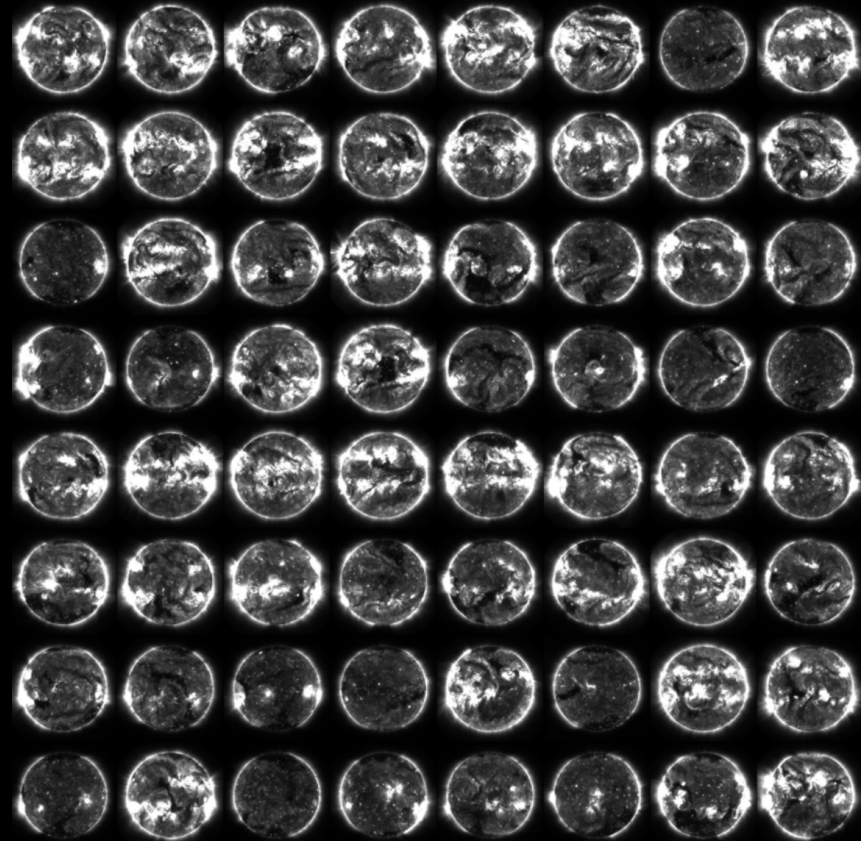
J. Zbinden



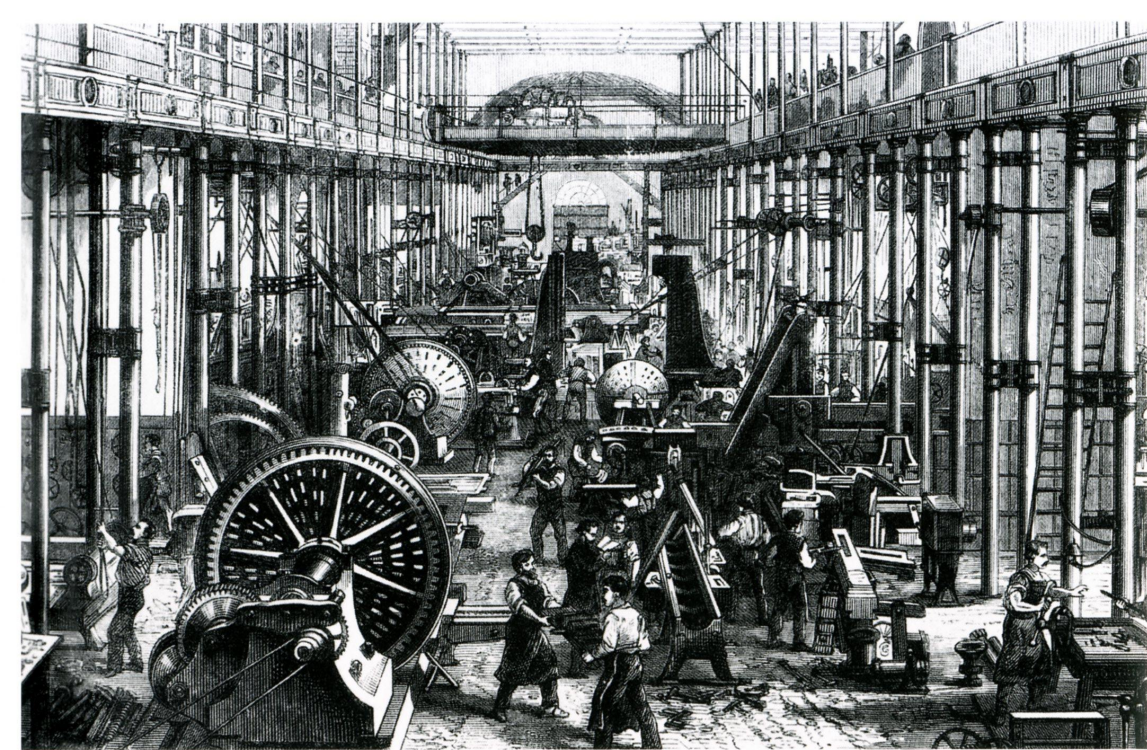
MNIST



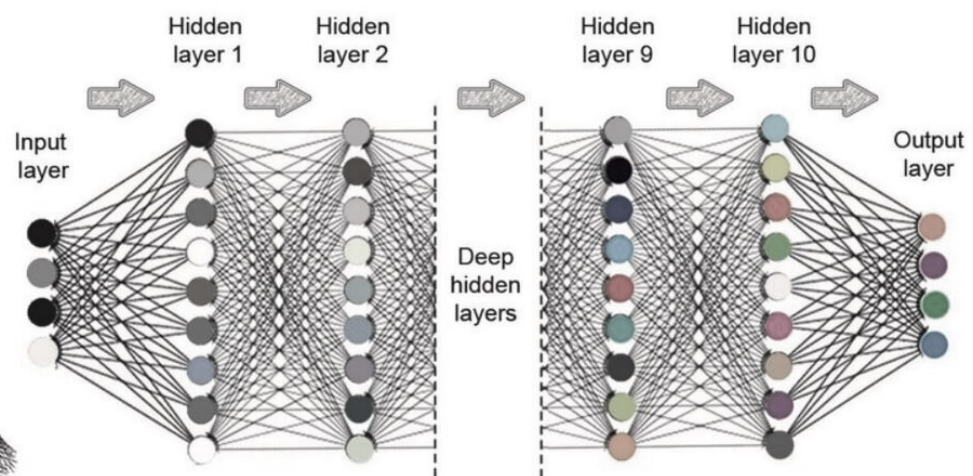
NASA/SDO



These techniques are data agnostic, so we can use them to solve our open questions

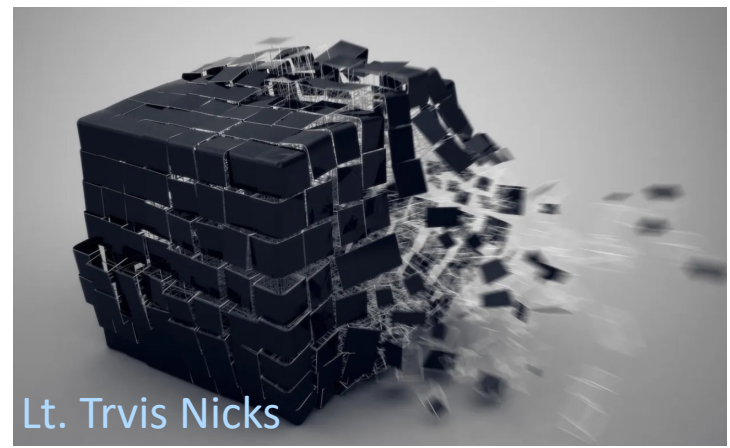


ML was nurtured by **industry**, where outcome supersedes understanding



P. G. Breen et.al MNRAS 2020, illustration by Mahala Le May

The **deep learning** revolution further reduced the role of understanding



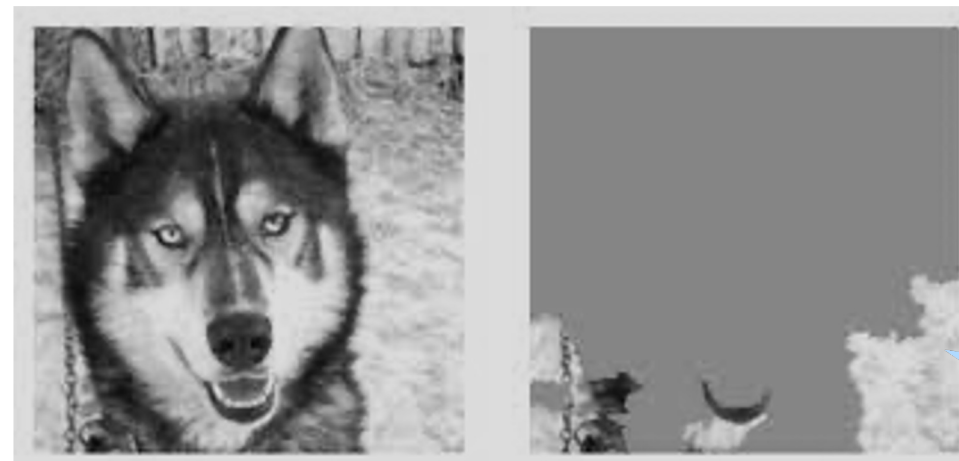
Lt. Trvis Nicks

Neural networks effectively function as efficient **black boxes**

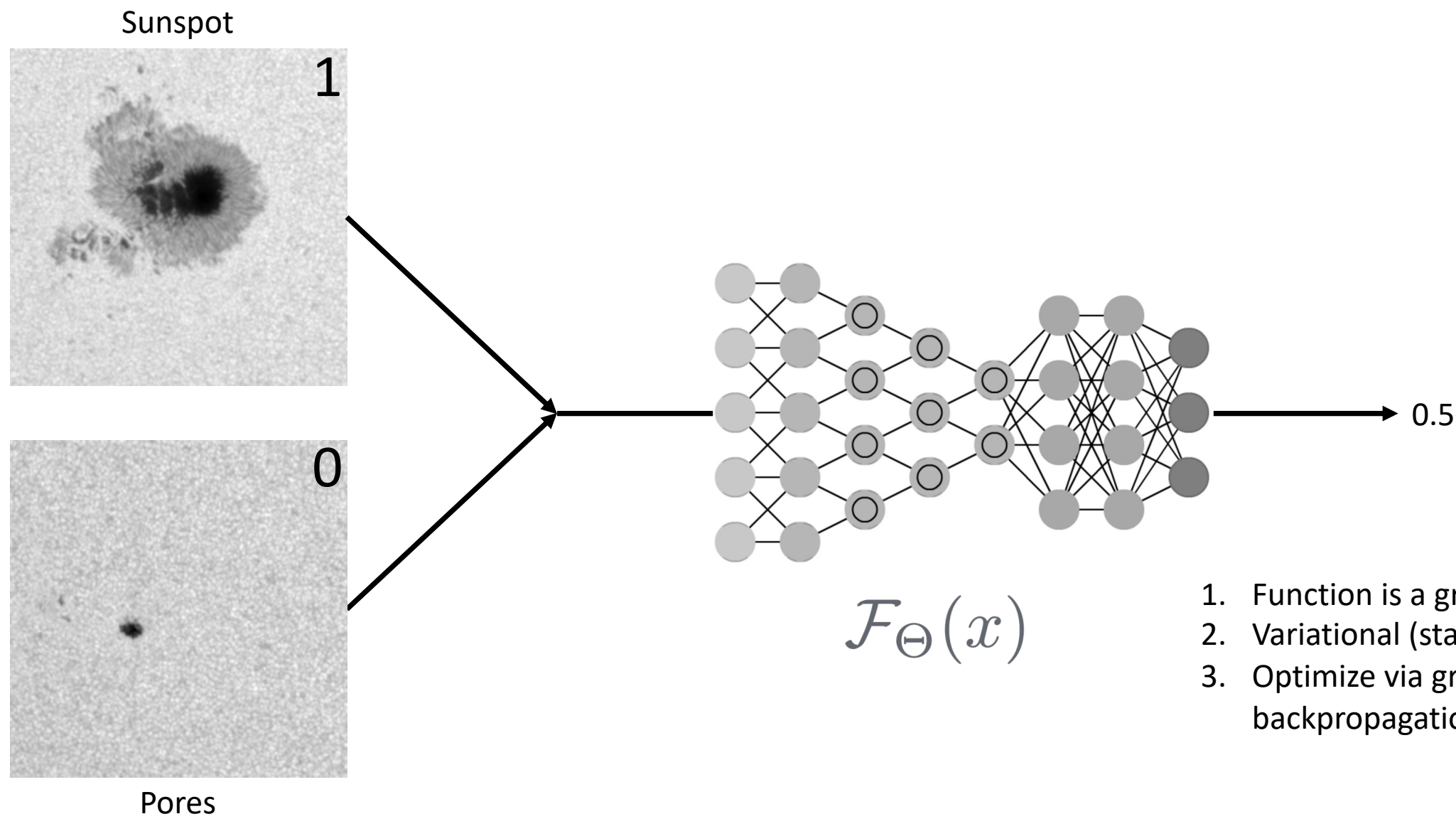
Ribeiro et. al, *Why Should I Trust You?* (2016)



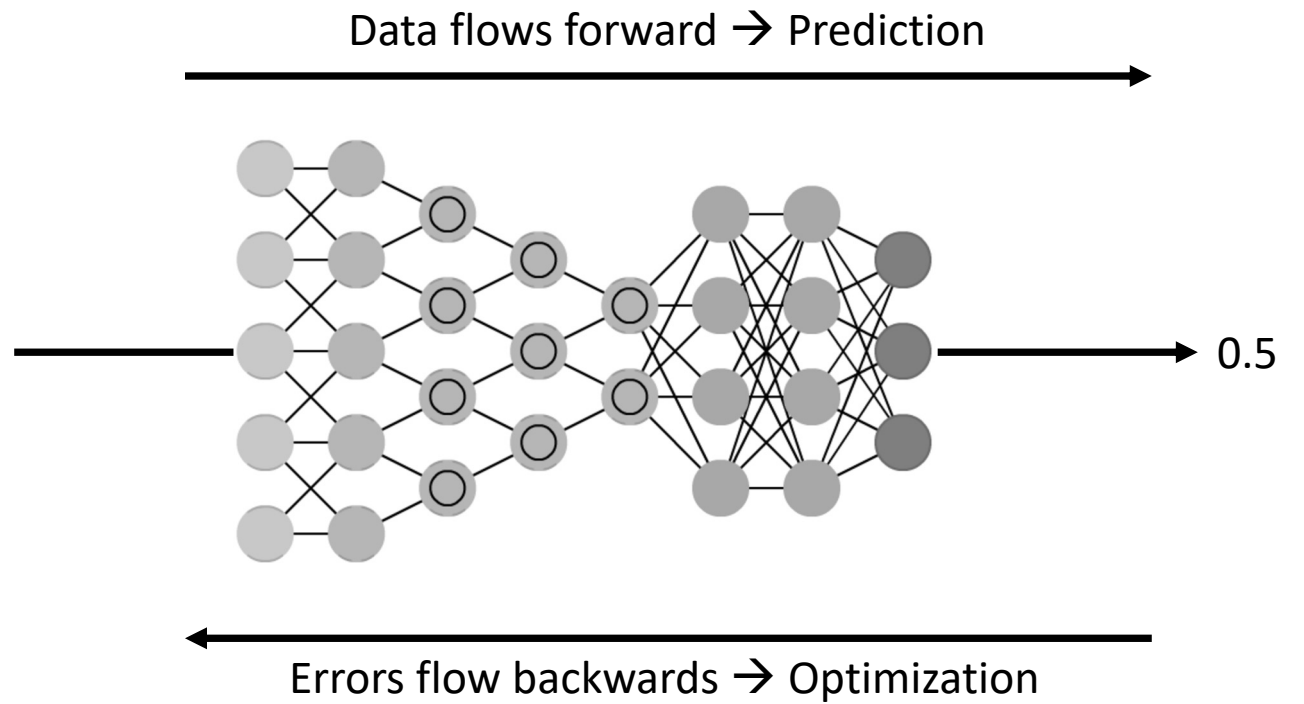
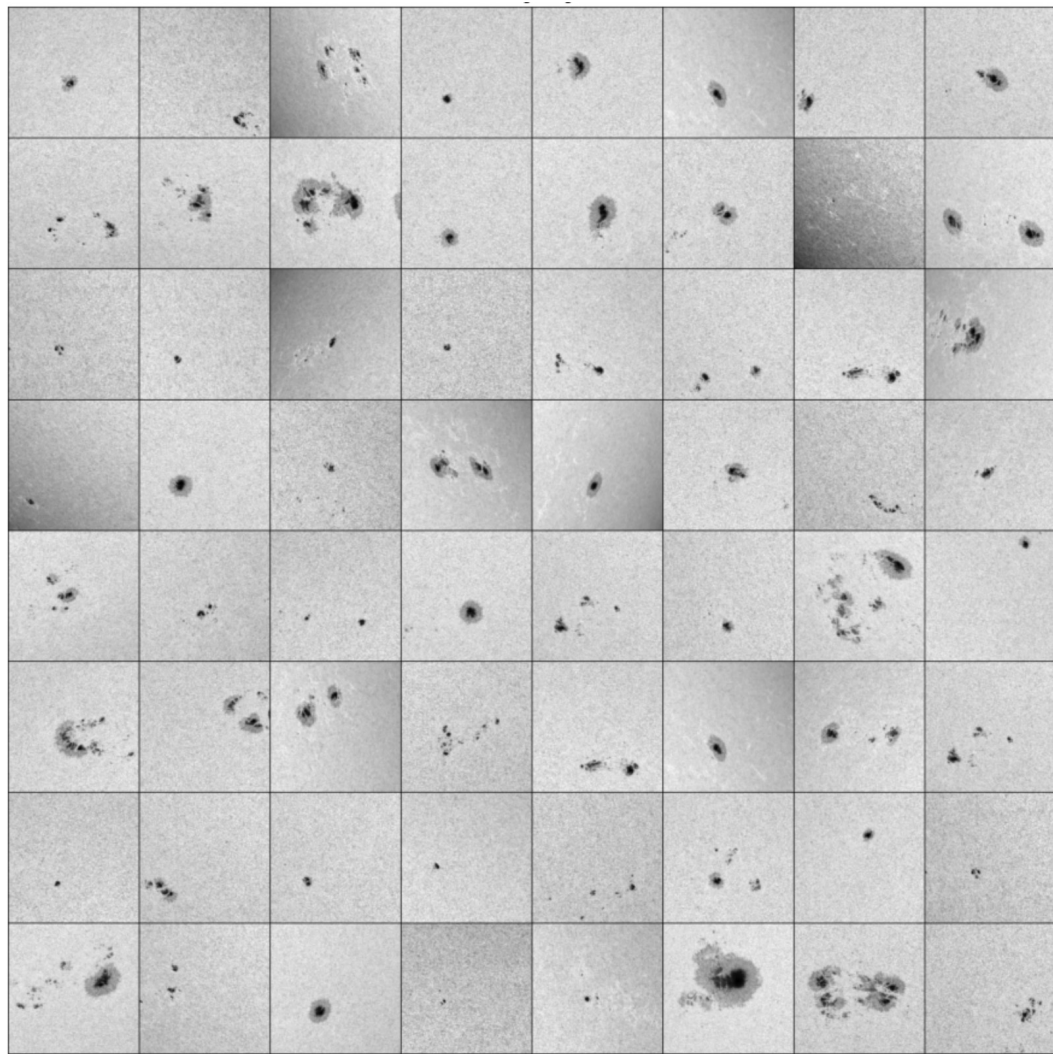
Ribeiro et. al, *Why Should I Trust You?* (2016)

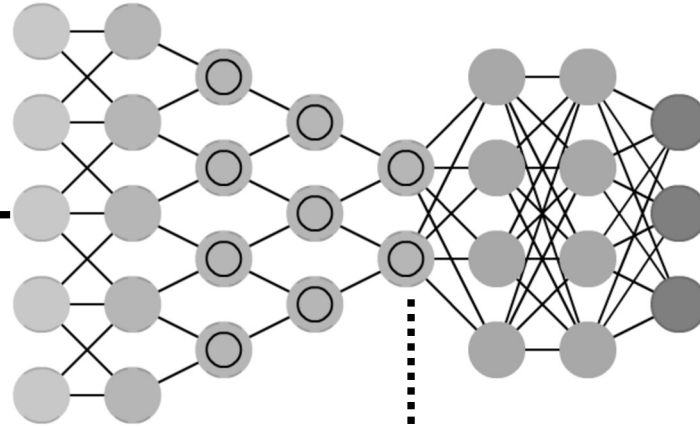
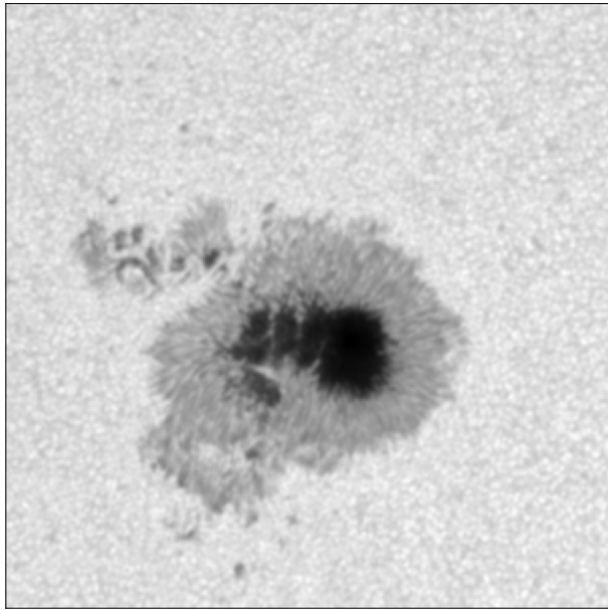


Snow



1. Function is a graph Like structure
2. Variational (start with bad guess)
3. Optimize via gradient descent using backpropagation

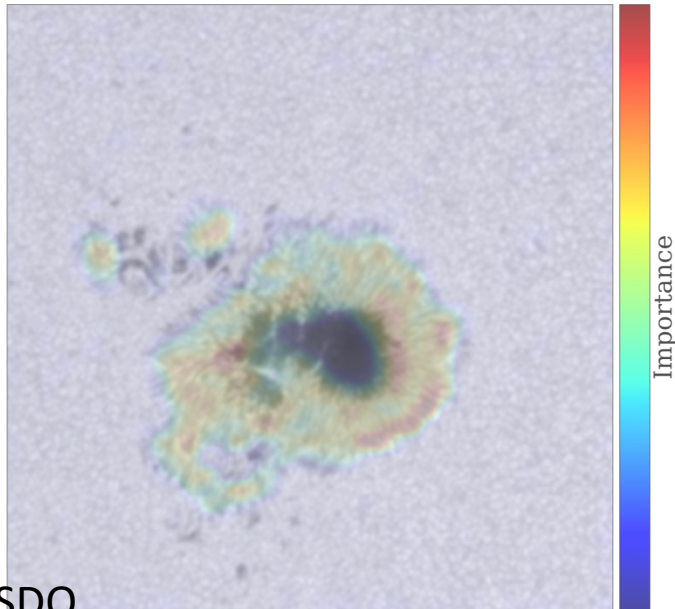




0.9

Explainable AI

Grad-CAM ([Selvaraju et al. 2017](#))



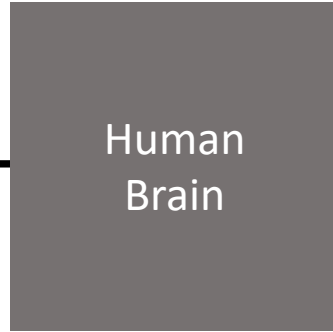
Grad-CAM allows us to see what distinguishes a sunspot from a pore

Penumbra!

Reality basis

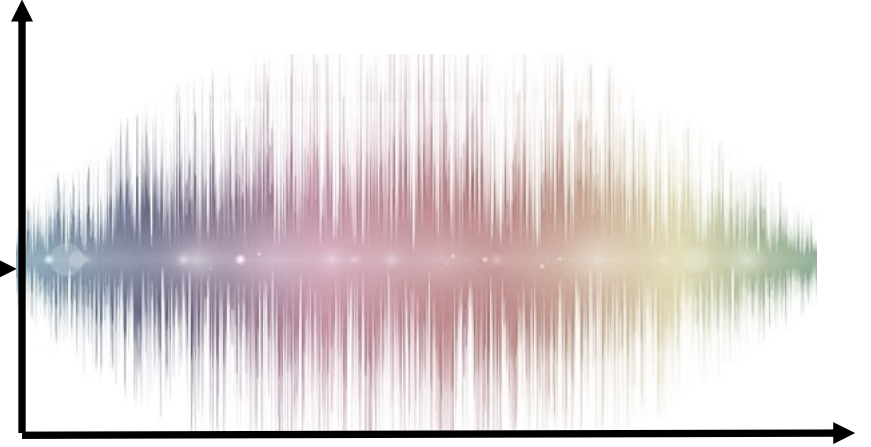
$$\begin{aligned} \frac{\partial}{\partial x_i} \frac{\partial}{\partial x_k} A_i - \frac{\partial}{\partial x_i} \frac{\partial}{\partial x_i} A_k + \frac{1}{c} \frac{\partial}{\partial x_k} \frac{\partial \phi}{\partial t} + \frac{1}{c^2} \frac{\partial^2 A_k}{\partial t^2} &= \frac{4\pi}{c} J_k \\ \frac{\partial}{\partial x_k} \vec{\nabla} \cdot \vec{A} - \nabla^2 A_k + \frac{1}{c} \frac{\partial}{\partial x_k} \frac{\partial \phi}{\partial t} + \frac{1}{c^2} \frac{\partial^2 A_k}{\partial t^2} &= \frac{4\pi}{c} J_k \\ -\nabla^2 A_k + \frac{1}{c^2} \frac{\partial^2 A_k}{\partial t^2} + \frac{\partial}{\partial x_k} \left(\vec{\nabla} \cdot \vec{A} + \frac{1}{c} \frac{\partial \phi}{\partial t} \right) &= \frac{4\pi}{c} J_k \\ -\nabla^2 \vec{A} + \frac{1}{c^2} \frac{\partial^2 \vec{A}}{\partial t^2} + \vec{\nabla} \left(\vec{\nabla} \cdot \vec{A} + \frac{1}{c} \frac{\partial \phi}{\partial t} \right) &= \frac{4\pi}{c} \vec{J} \end{aligned}$$

Maxwell's equations

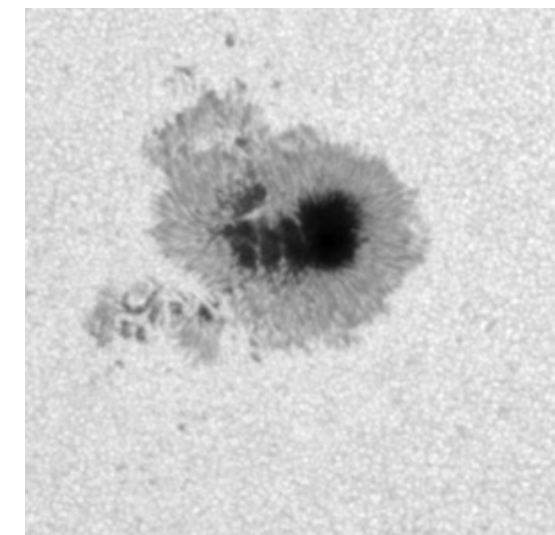


Abstraction

Human basis



More useful basis of colors



NASA/SDO

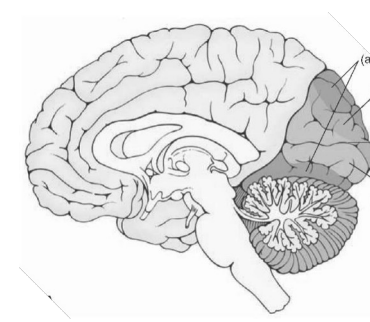
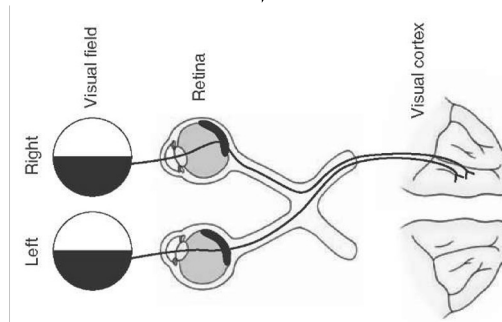
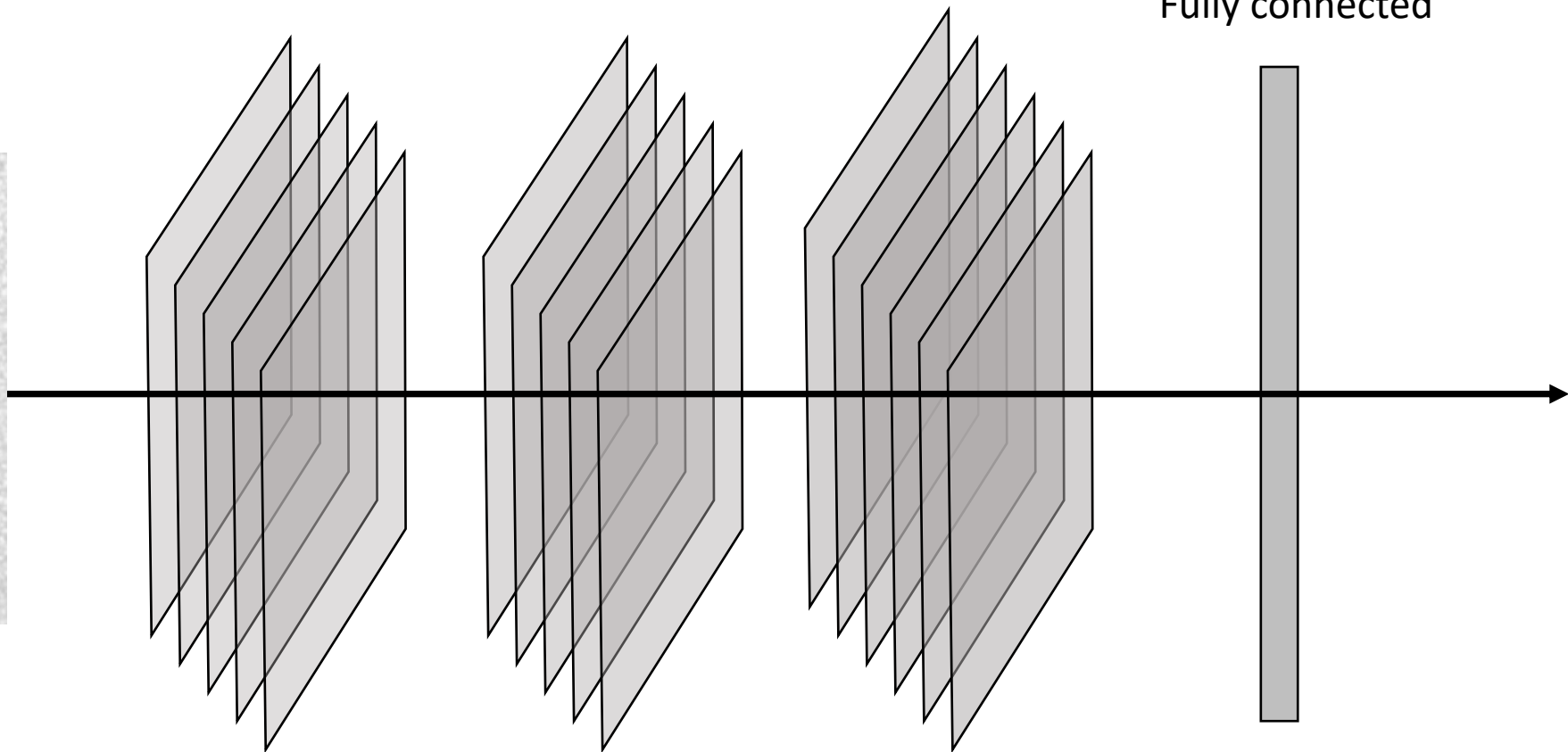
ConvLayer1

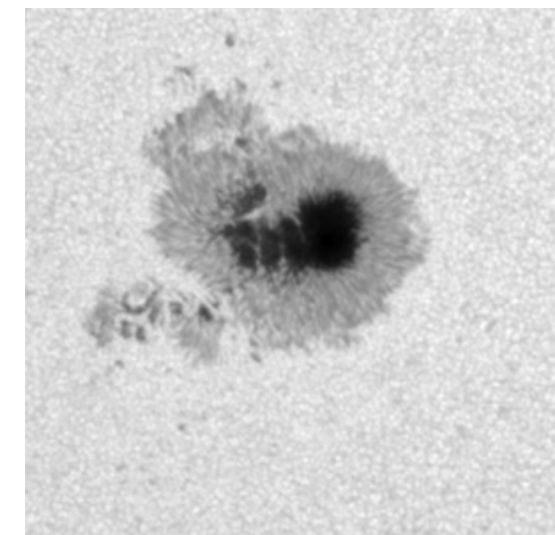
ConvLayer2

ConvLayer3

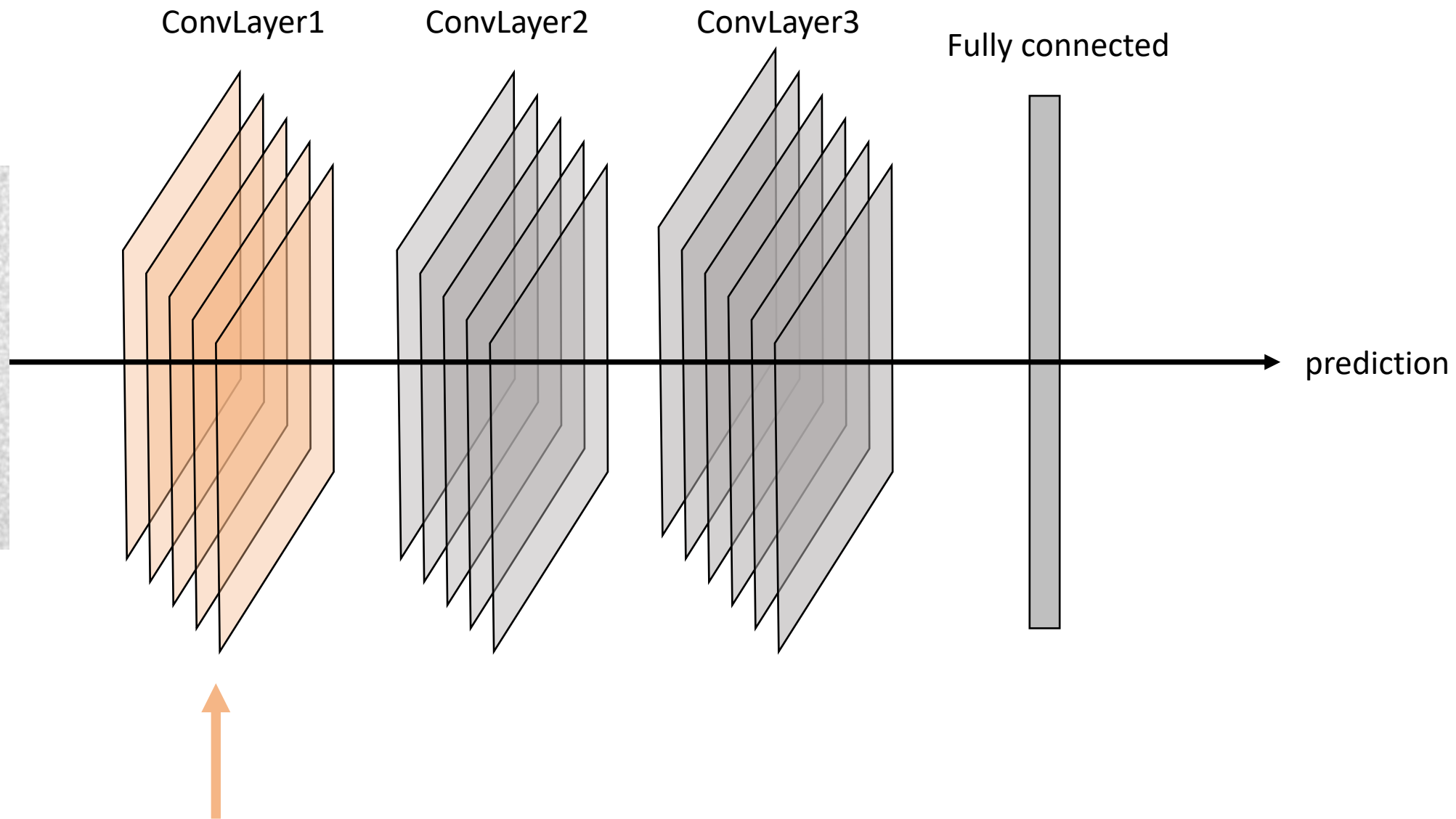
Fully connected

prediction

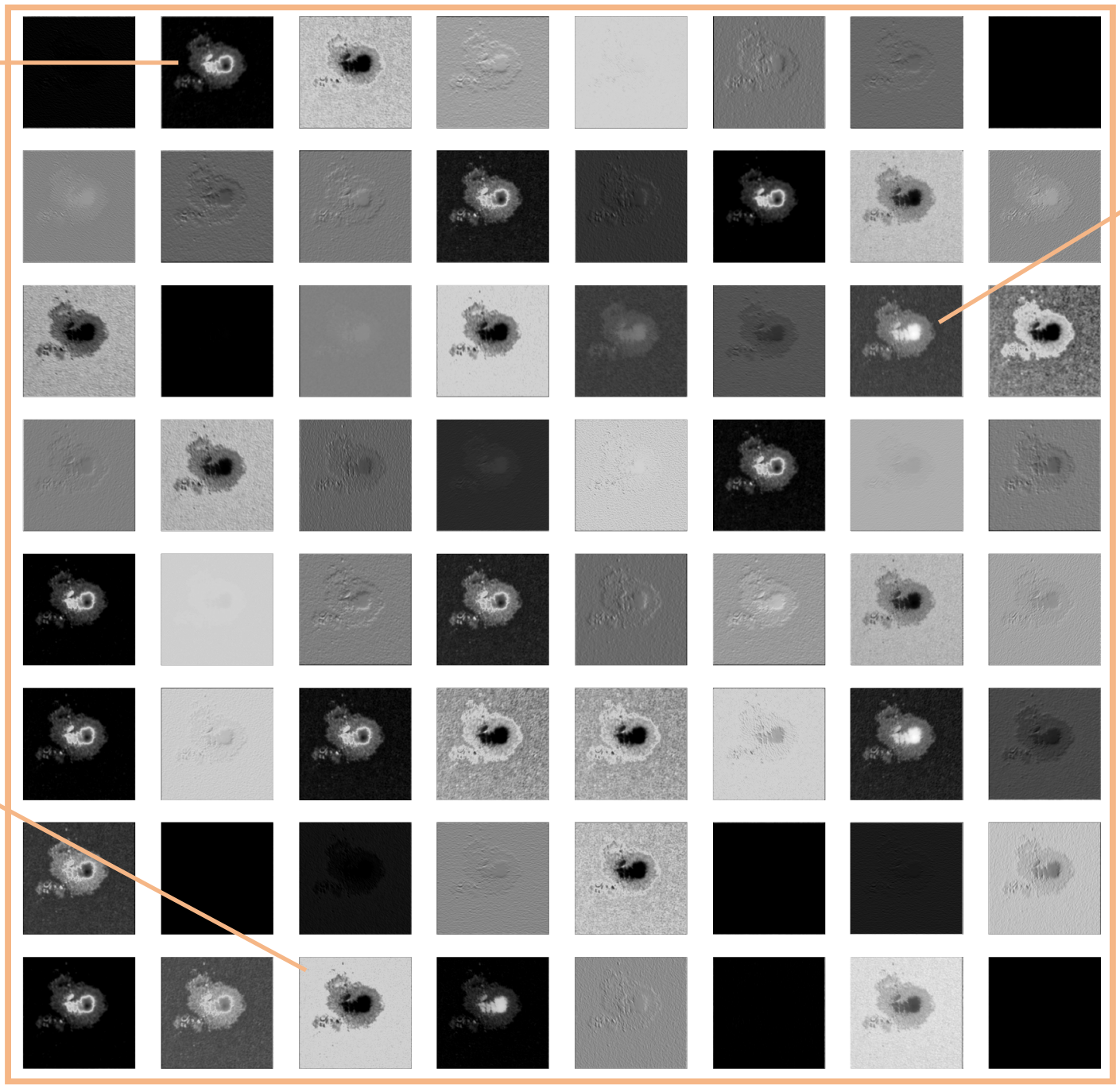




NASA/SDO



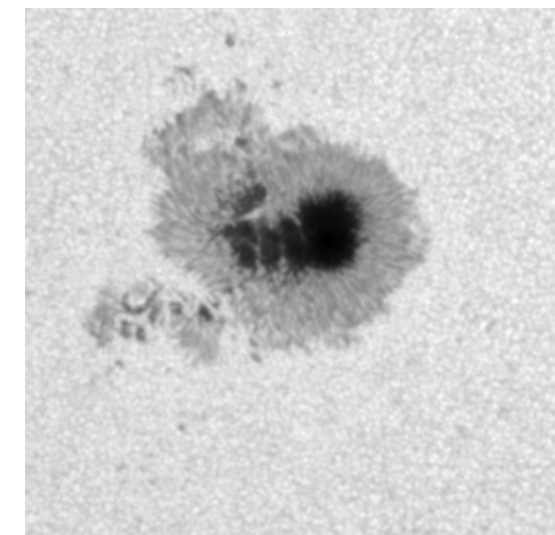
Penumbral border



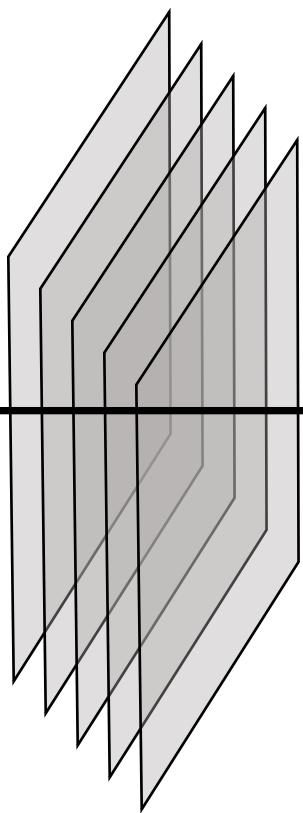
Umbra

Attention

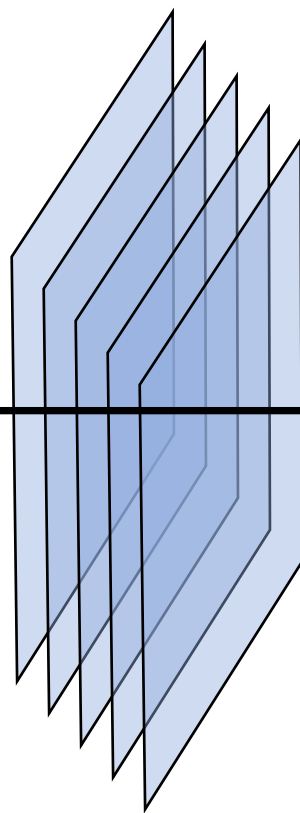
Granules



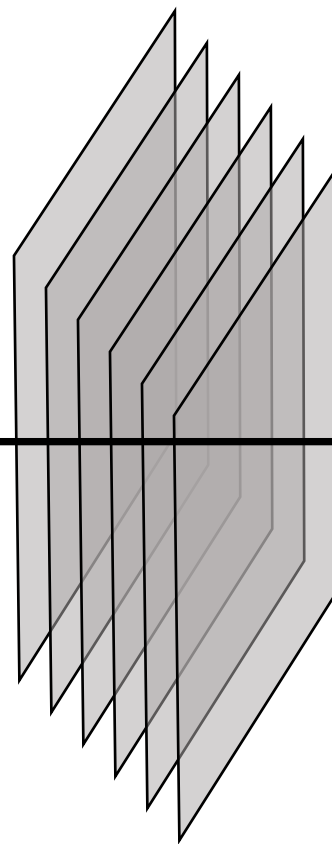
ConvLayer1



ConvLayer2



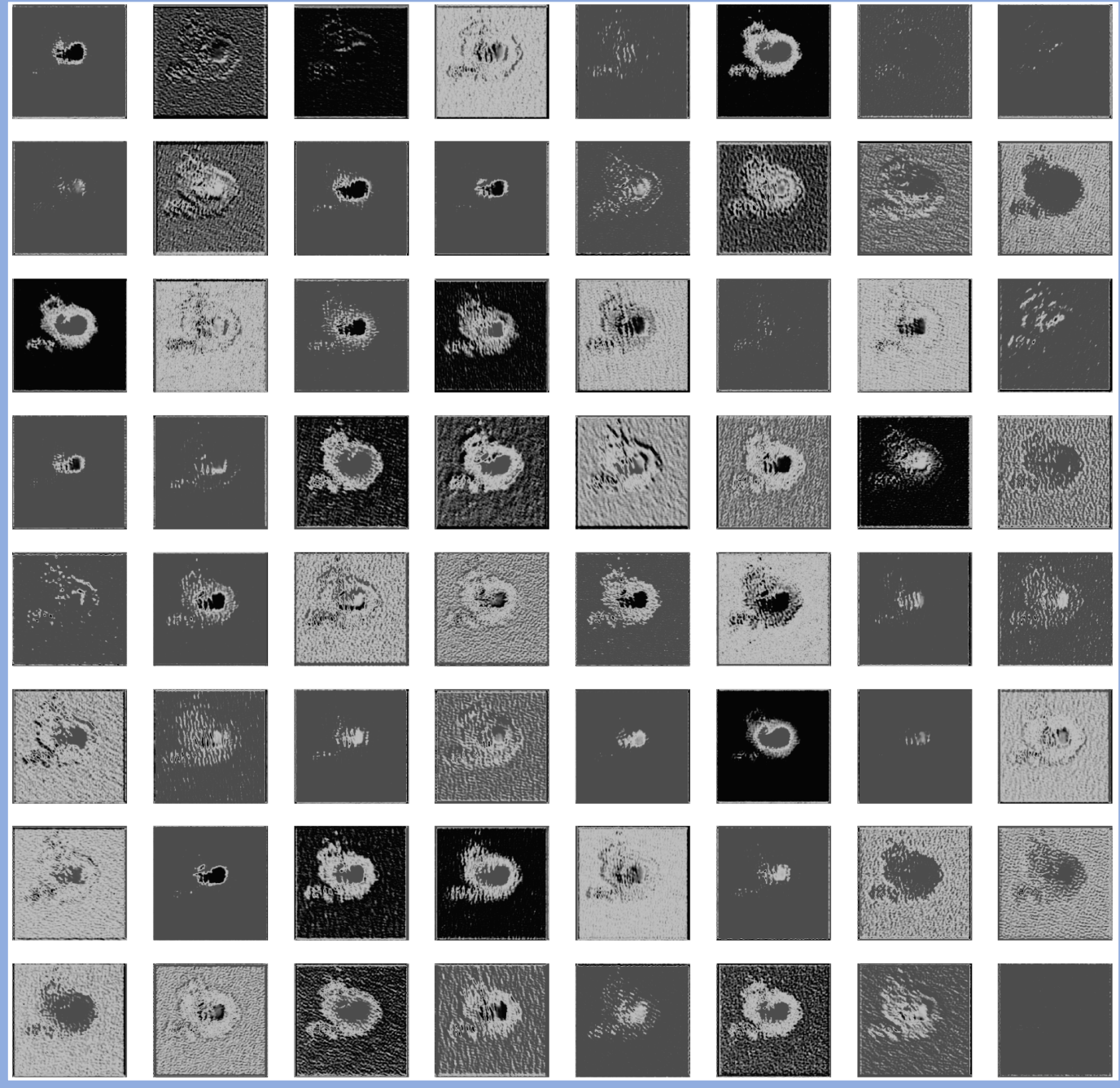
ConvLayer3



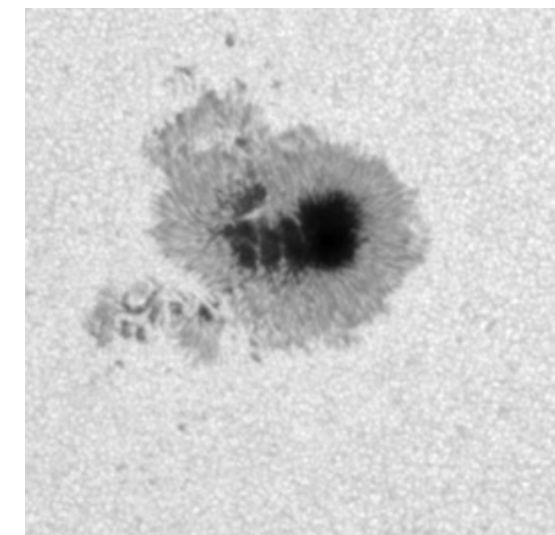
Fully connected



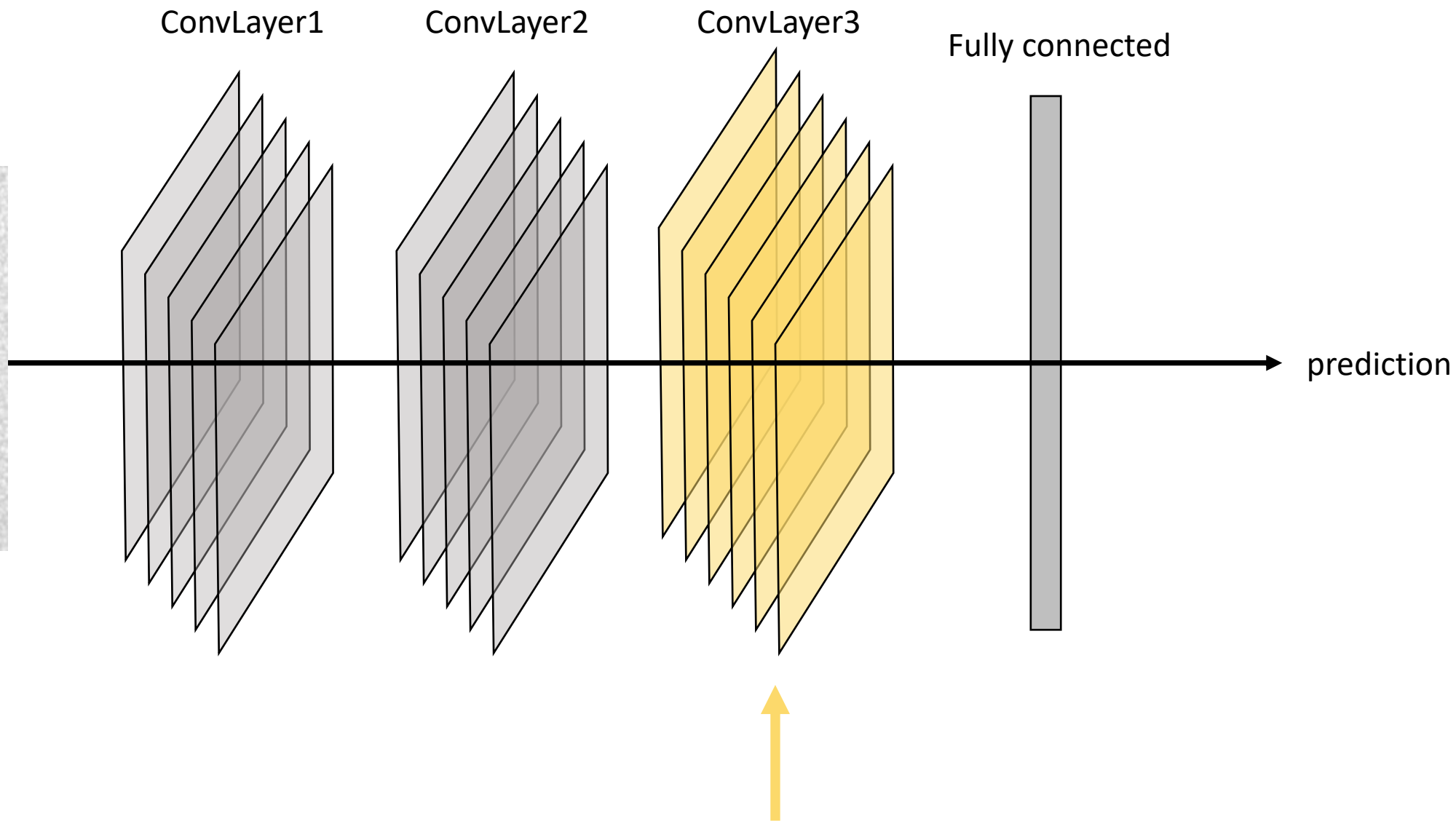
prediction



Attention

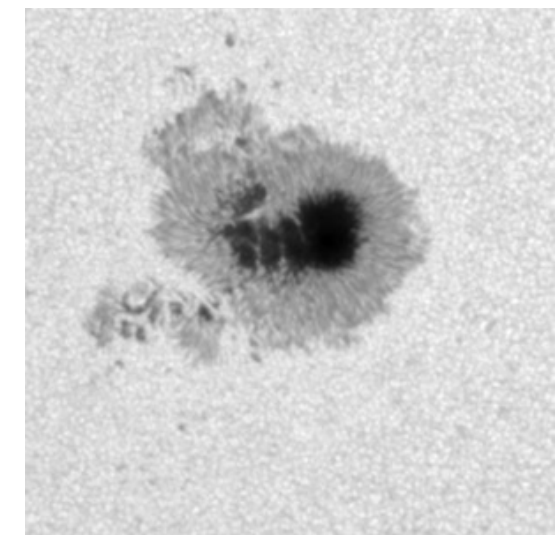


NASA/SDO

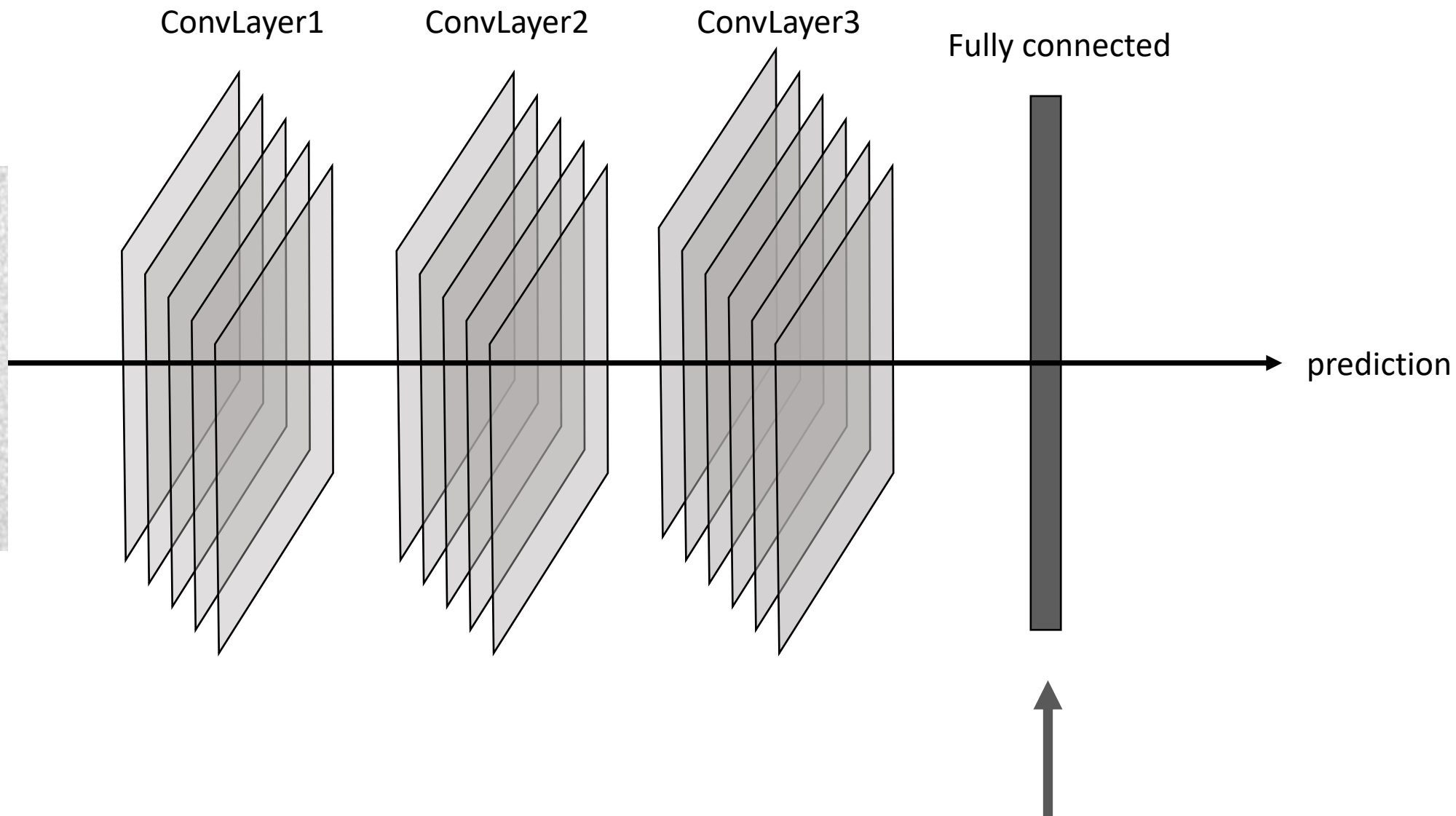




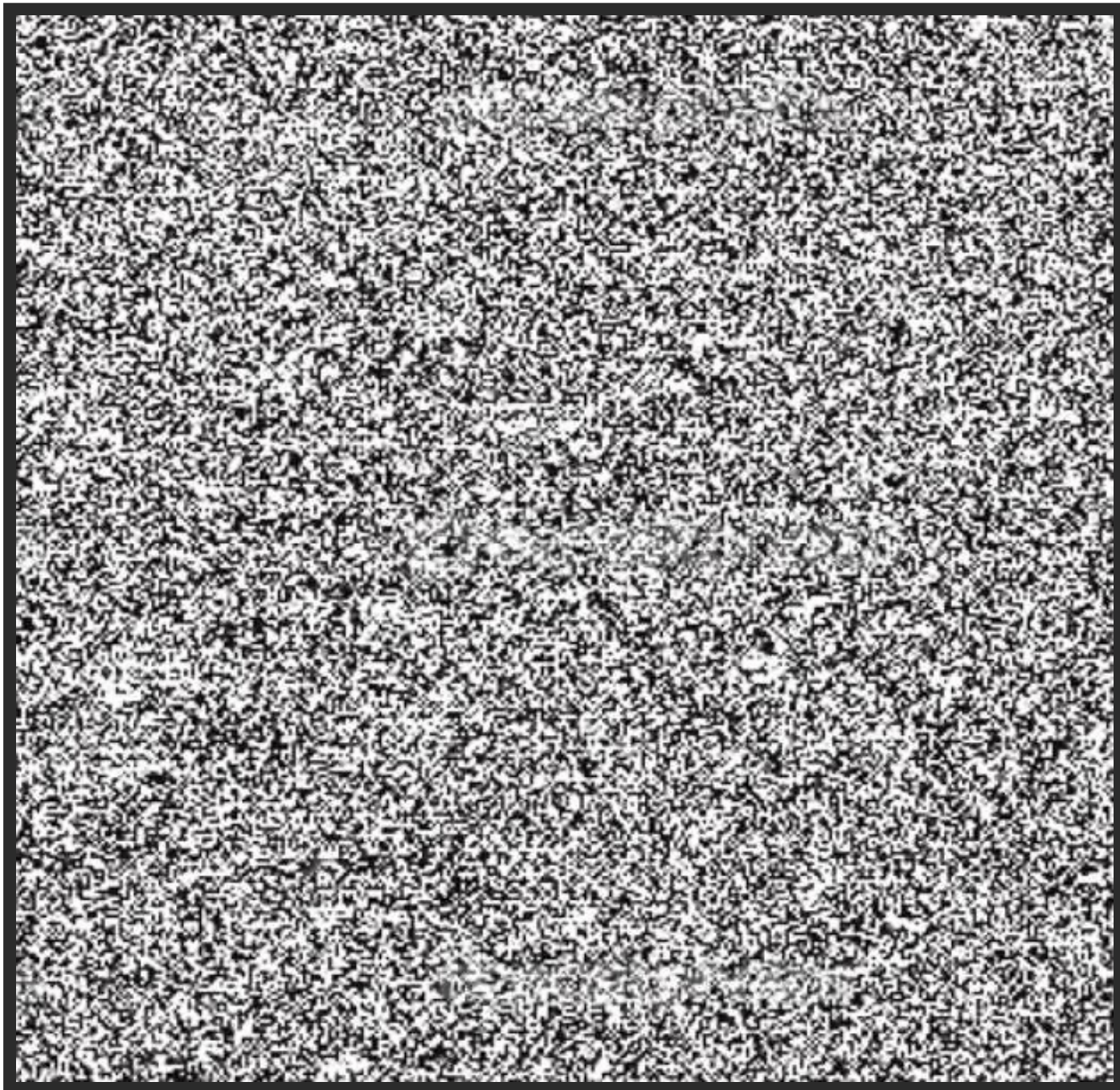
Attention



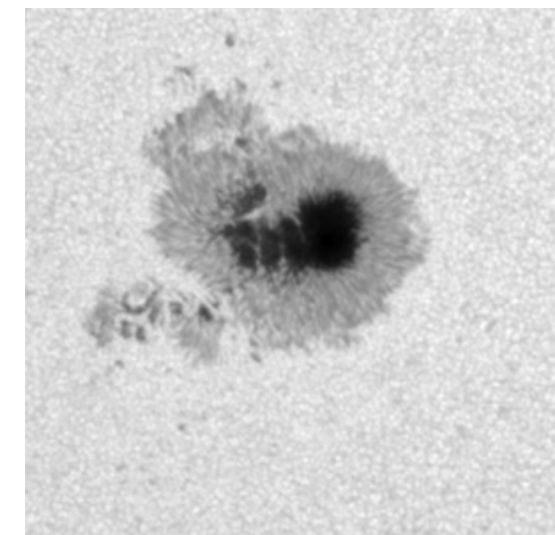
NASA/SDO



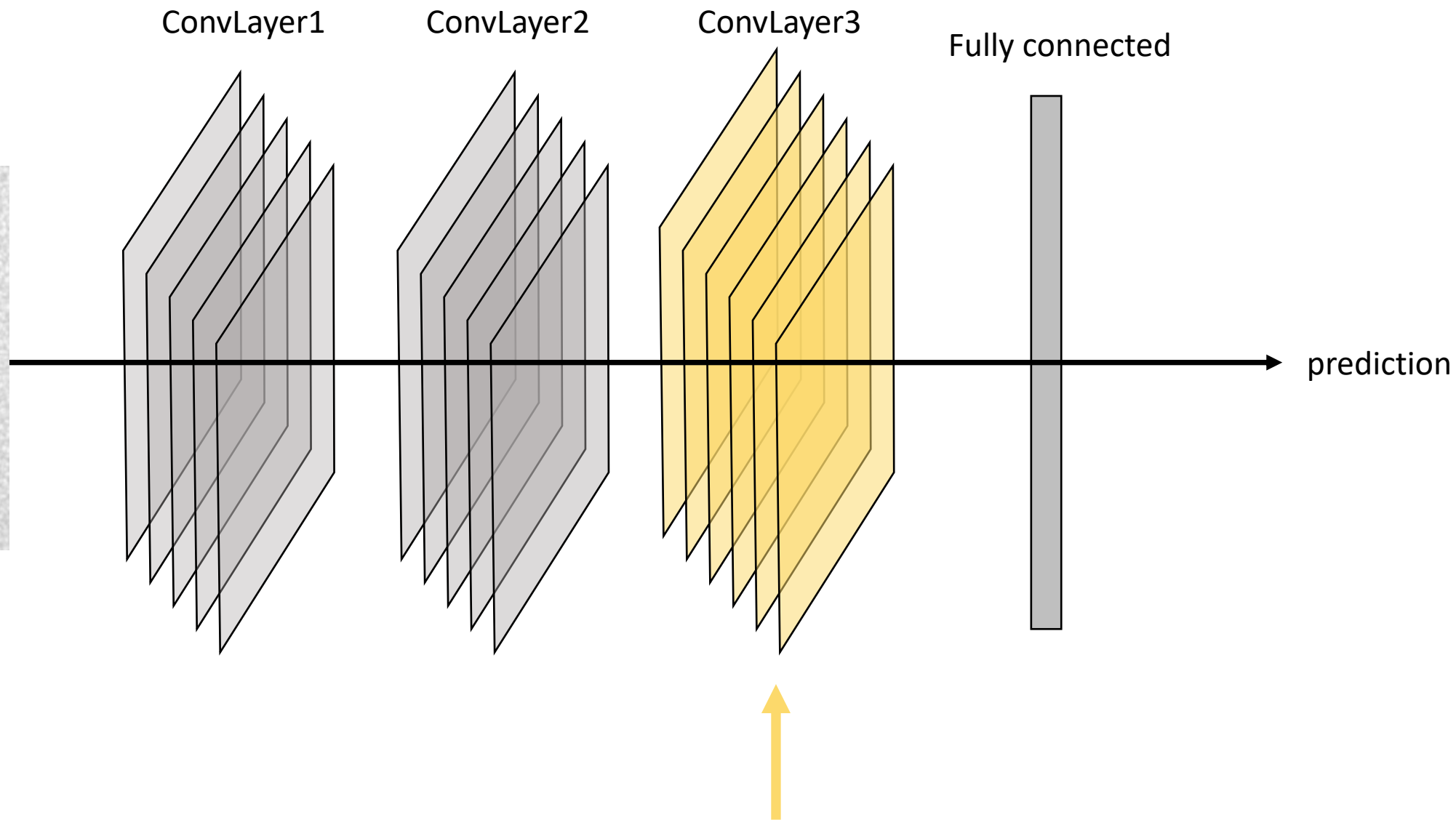
White Noise
Spatial coherence Is lost



Attention



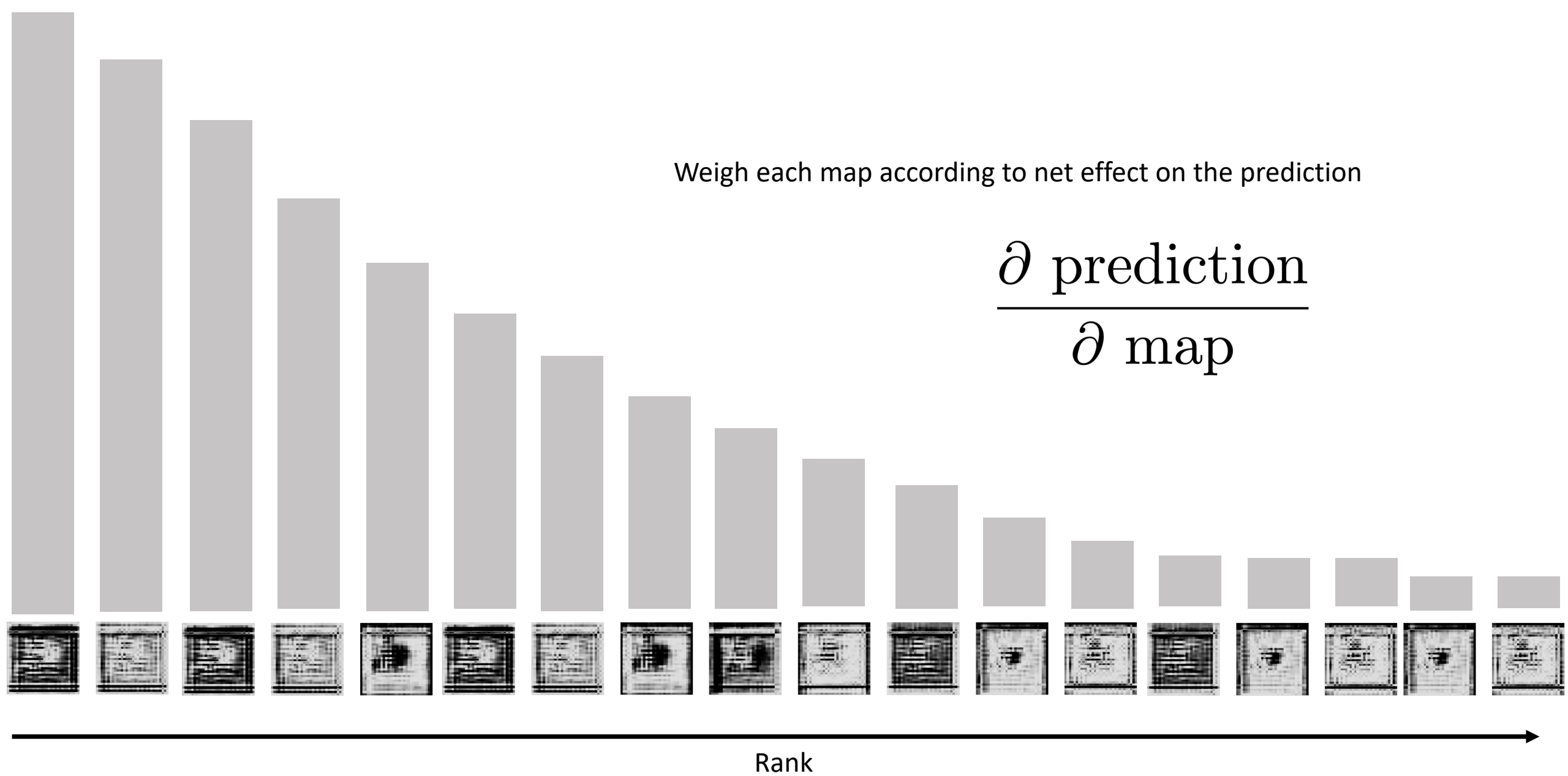
NASA/SDO



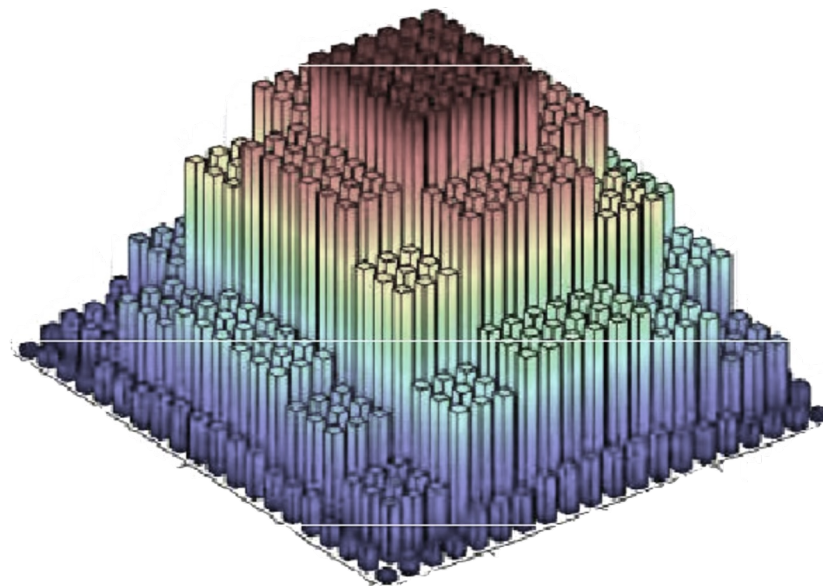
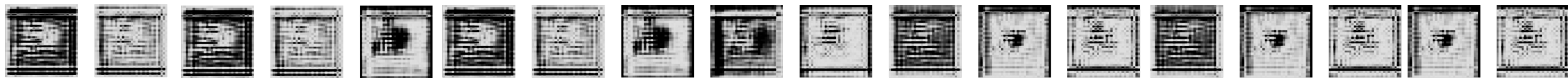
Wiggle each map and see how sensitive the output score is



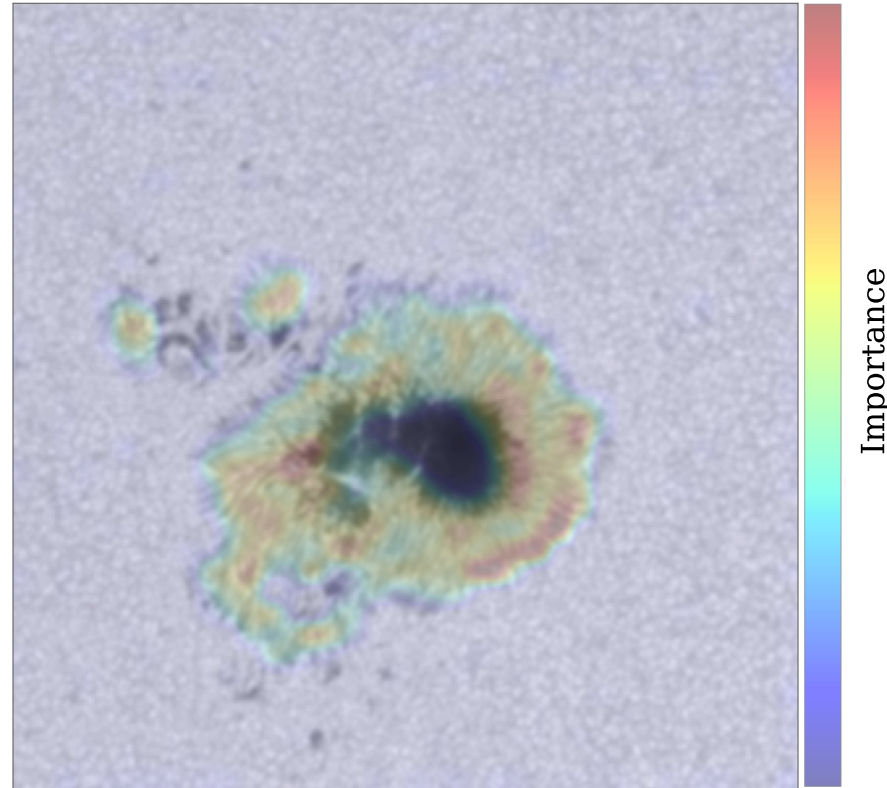
$$\frac{\partial \text{prediction}}{\partial \text{map}}$$



$$\sum_i \lambda_i \text{map}_i$$



The height (z) dimension is pixel importance



Grad-CAM highlights the penumbra, since this is what it searches for when classifying an image as a Sunspot

Blackbox model

Input datapoint

$$\phi_i(f, x) = \sum_{z' \subseteq x'} \frac{|z'|!(M - |z'| - 1)!}{M!} [f_x(z') - f_x(z' \setminus i)]$$

Shapley value for feature i

Subset

Simplified data input

Weighting

Contribution



F/ Blackbox = Market → = Deep learning model

X = Coalition → image pixels

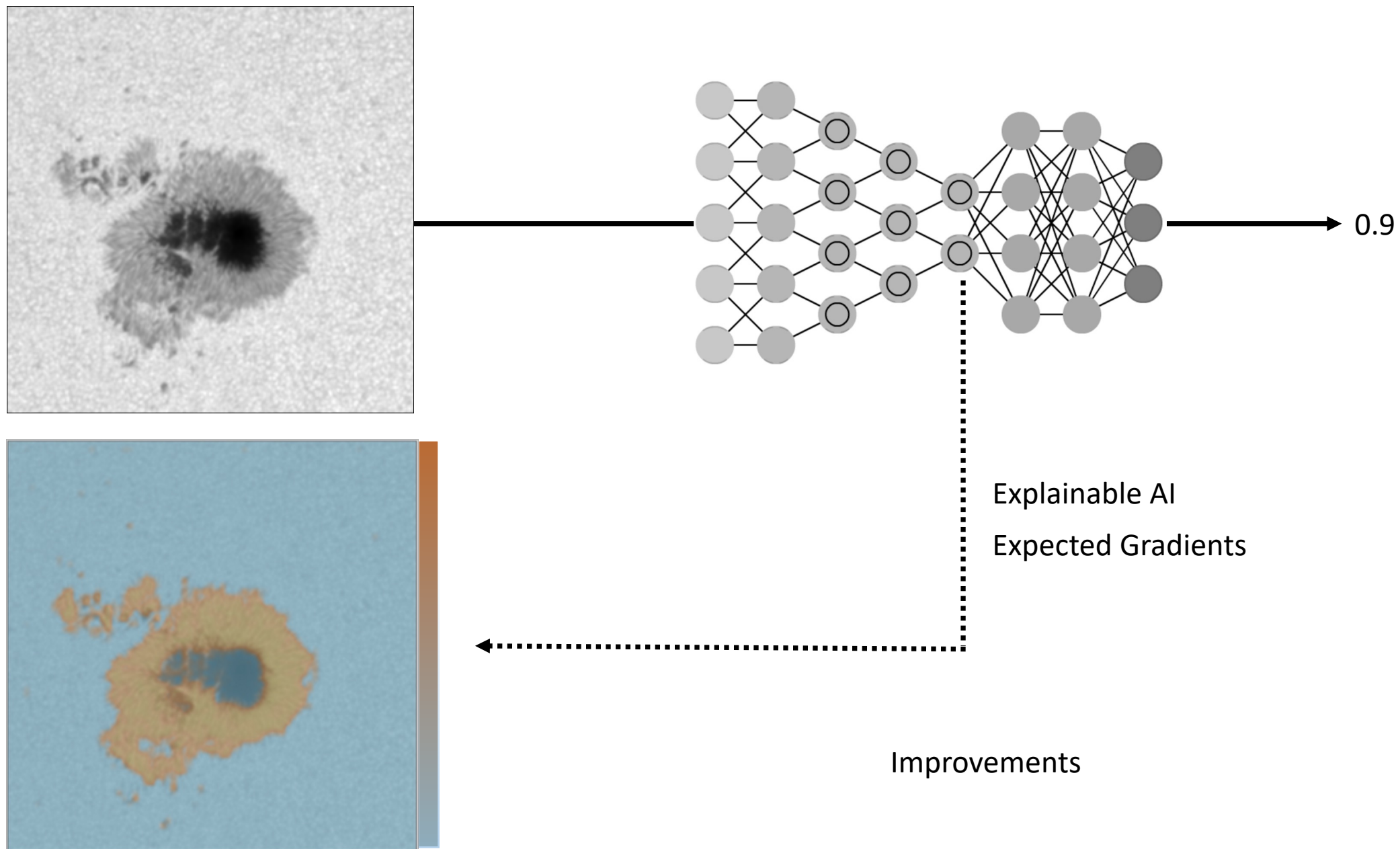
F(x) = pot of gold → prediction score [0,1]

Subset.....?

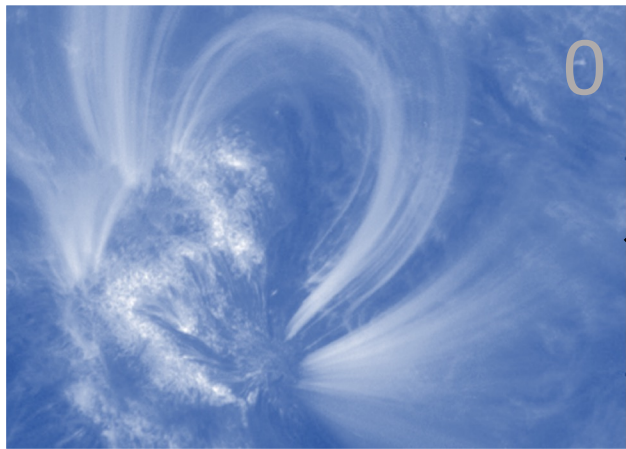
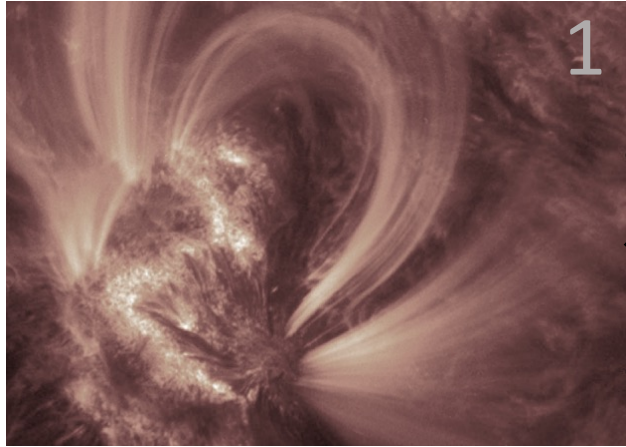
Zeros?, Gaussian?, Actual data?

$$\phi_{\lambda|EG} = \int_{x'} ((x_{\lambda} - x'_{\lambda}) \times \int_{\alpha=0}^1 \frac{\delta \mathcal{F}_{\Theta}(x' + \alpha(x - x'))}{\delta x_{\lambda}} d\alpha) p_D(x') dx'$$

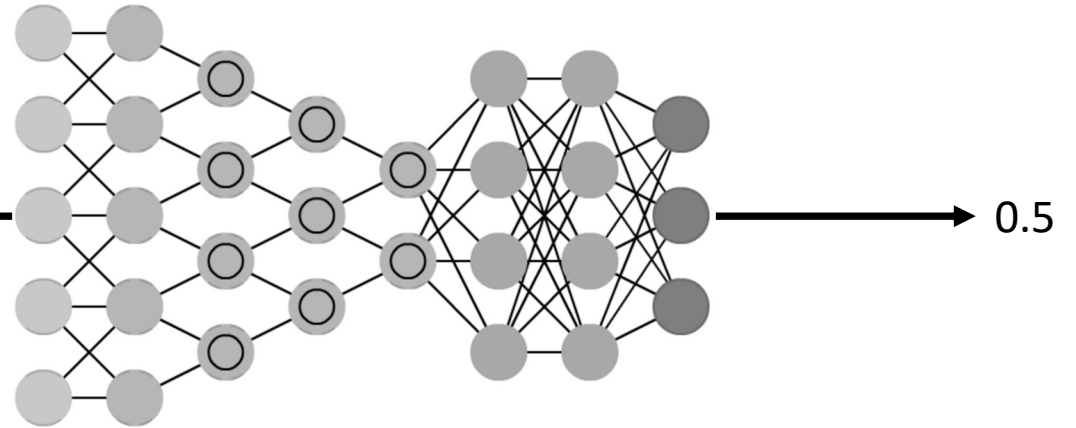
$$\approx \mathbb{E}_{x' \sim D, \alpha \sim U(0,1)} \left[(x_{\lambda} - x'_{\lambda}) \frac{\delta \mathcal{F}_{\Theta}(x' + \alpha \times (x - x'))}{\delta x_{\lambda}} \right],$$



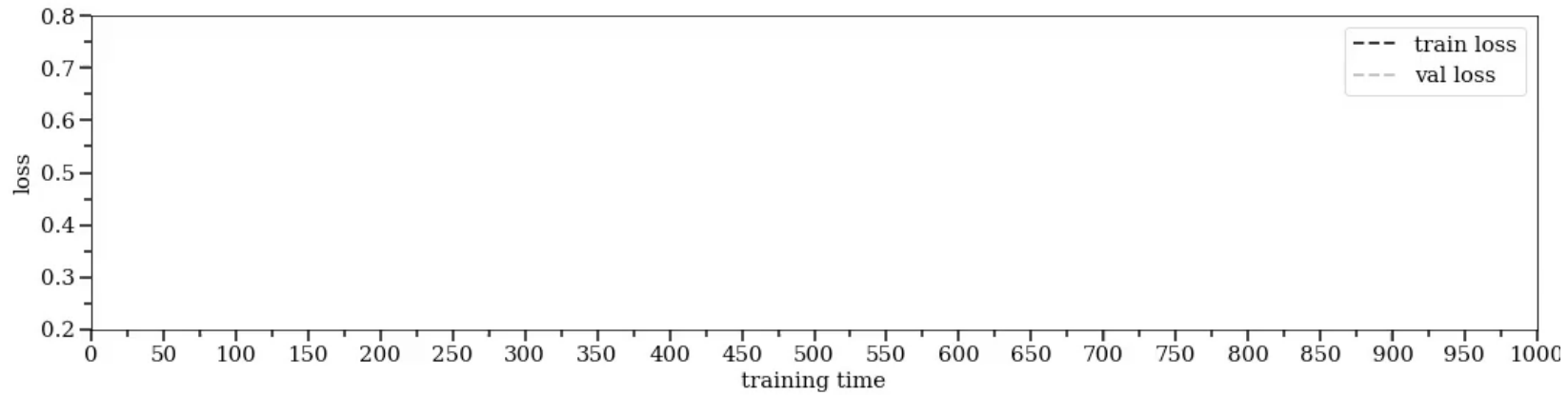
Spectra from pre-flare



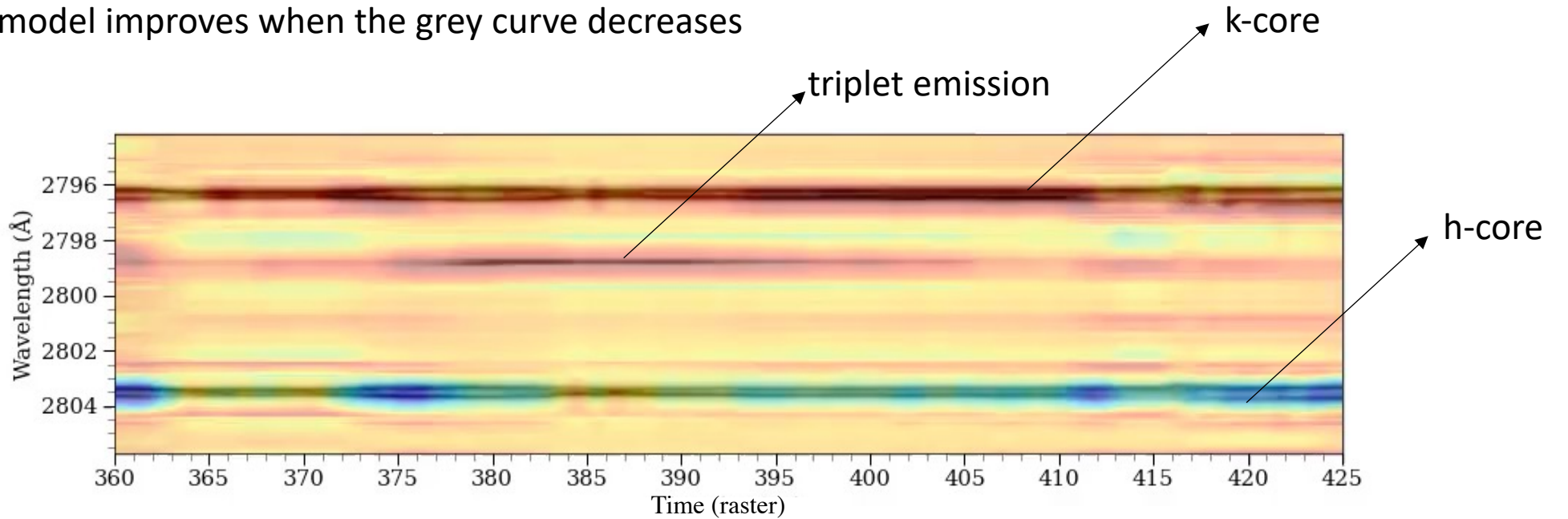
Spectra from active region



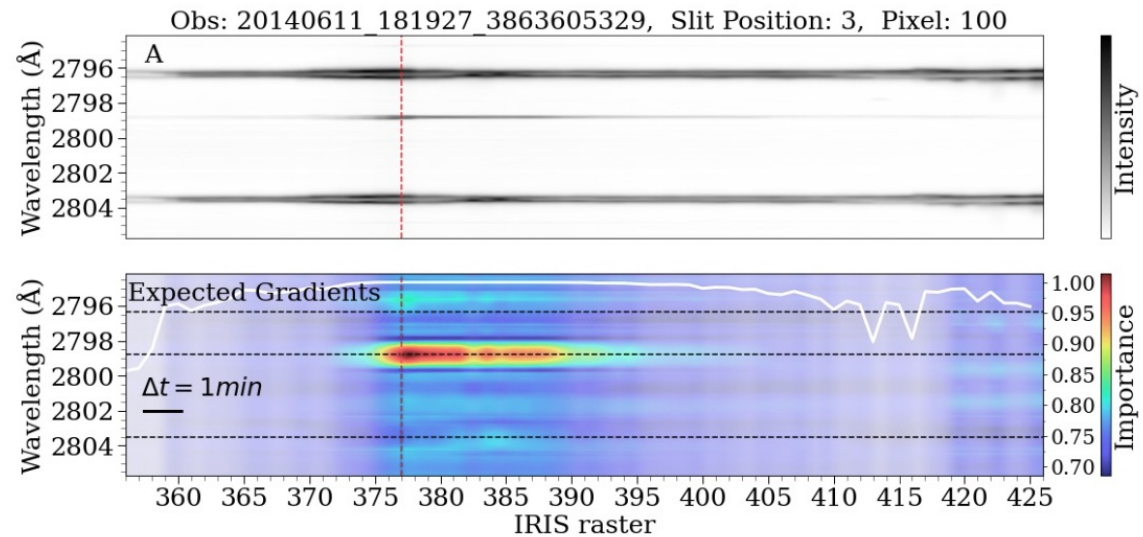
Once trained, the network should identify what features are strongly associated with precursory flare activity.



The model improves when the grey curve decreases

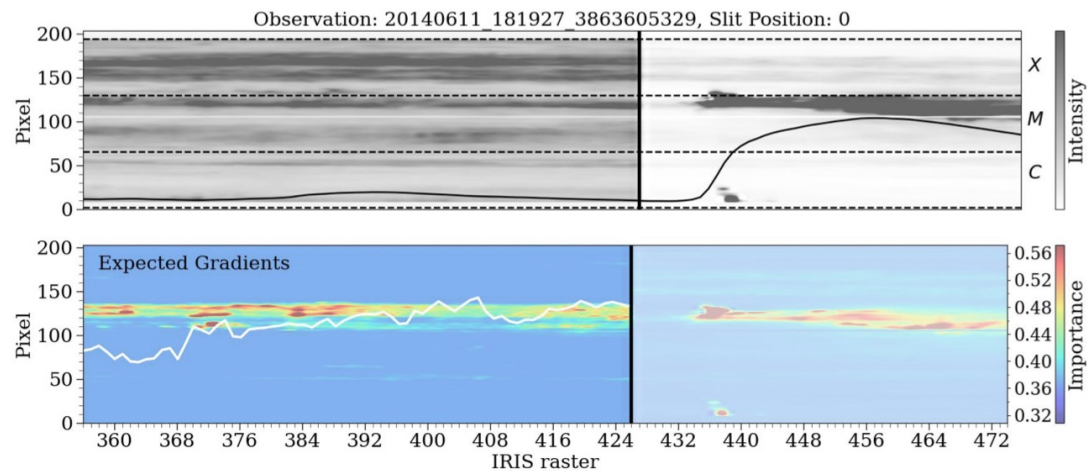
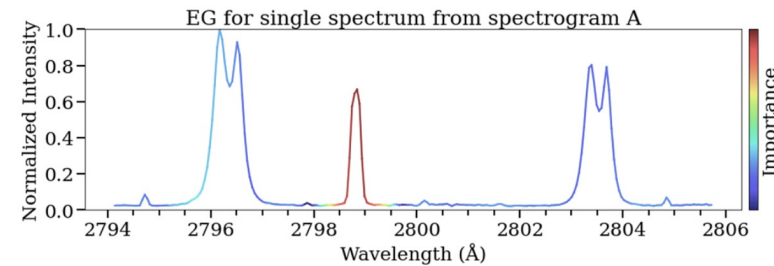


The network learns that triplet emission is important for deciding whether an active region will flare or not



B. Panos, L. Kleint, J. Zbinden A&A 671, A73 (2023)

Predicts triplet emission is nb for flare forecasting



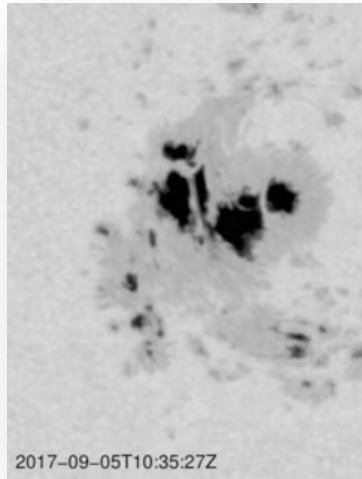
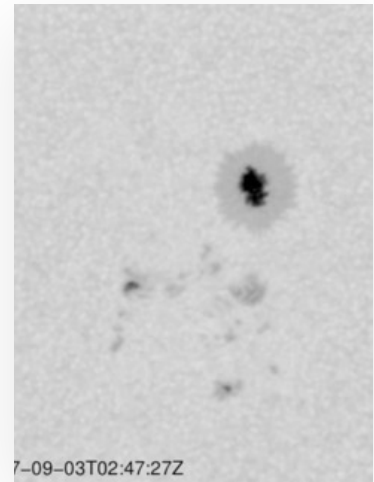
Model places attention where max future flare UV emission is seen later, even if more intense somewhere else

Modeling Sunspot evolution with transformer networks (Janis Witmer)

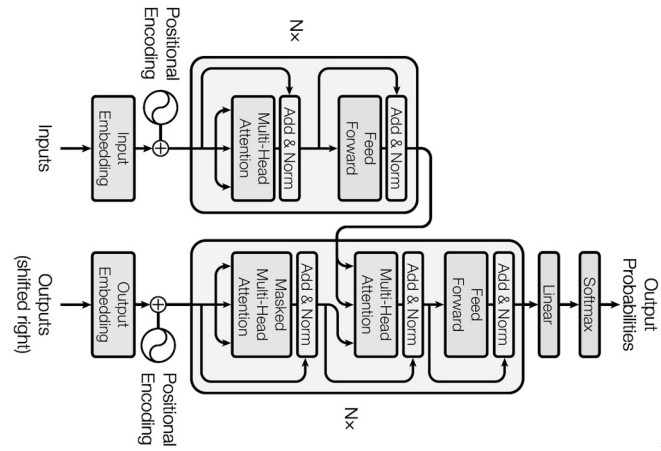


NASA/SDO

The idea

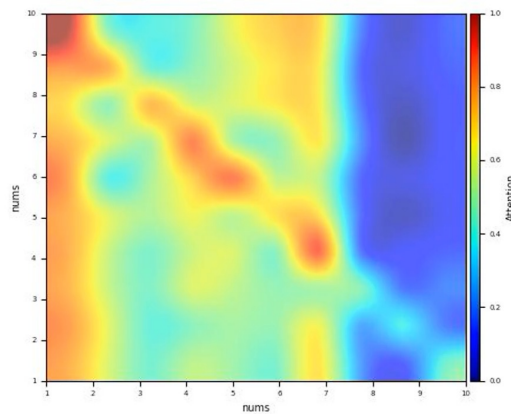


The model (ChatGPT)

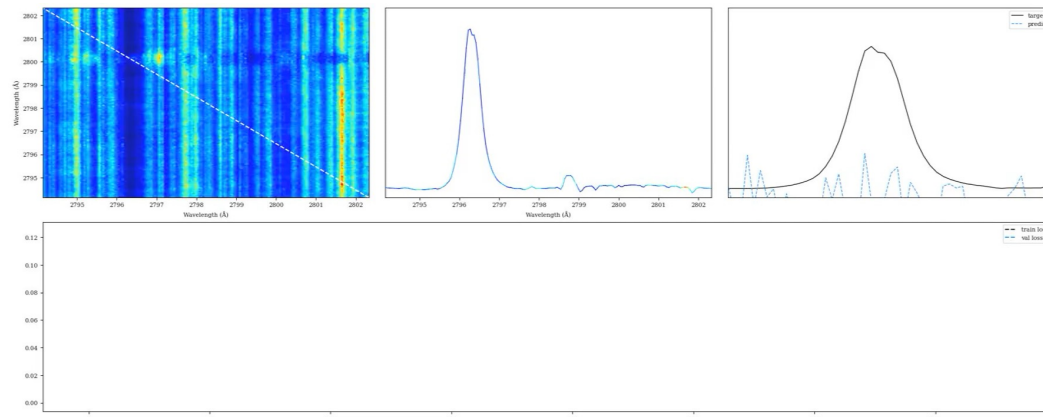


A. Vaswani et.al neurips (2017)

Reverse dataset



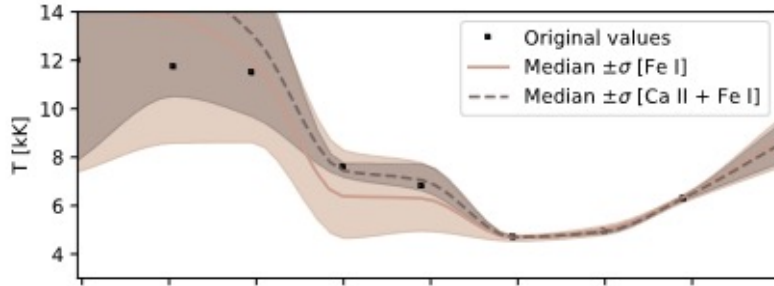
Spectral dataset



The tests

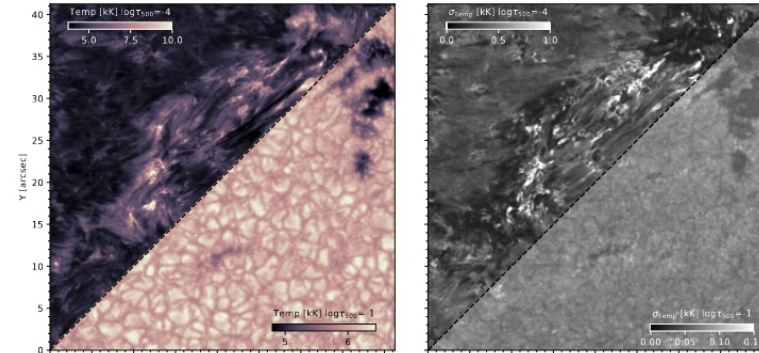
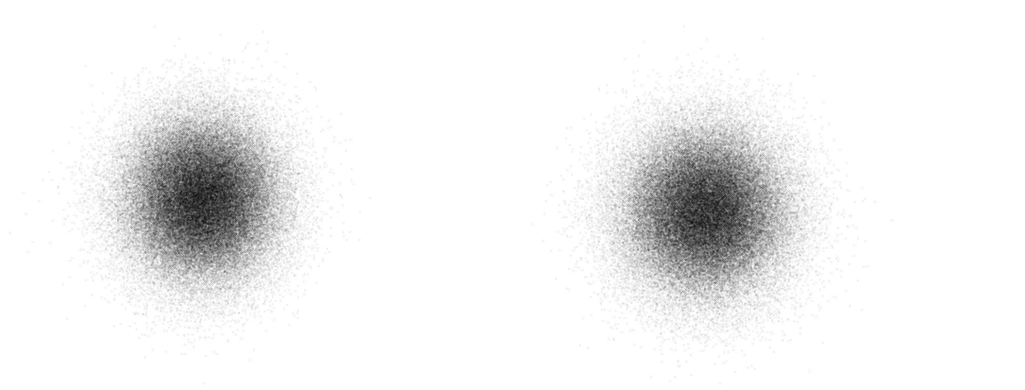
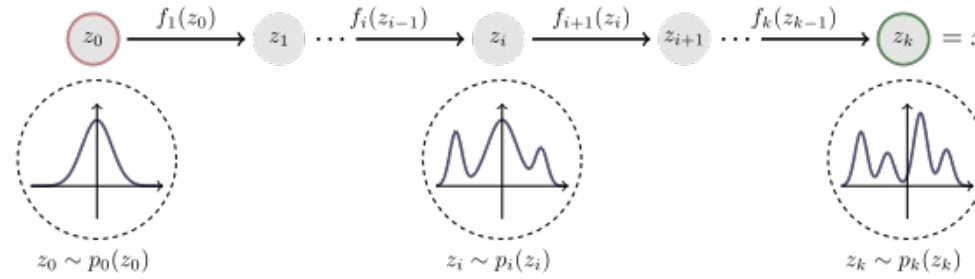


The idea



D. Baso et al A&A (2022)

The model (Normalizing flows)



D. Baso et al A&A (2022)

Thank you for your attention